Asteroid explorer, Hayabusa2, reporter briefing

July 25, 2017 JAXA Hayabusa2 Project



Topics



Regarding Hayabusa2,

• Results from the 2nd touchdown operation



Contents



- 0. Hayabusa2 and mission flow outline
- 1. Current status and overall schedule of the project
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Overview of Hayabusa2



Objective

We will explore and sample the C-type asteroid Ryugu, which is a more primitive type than the S-type asteroid Itokawa that Hayabusa explored, and elucidate interactions between minerals, water, and organic matter in the primitive solar system. By doing so, we will learn about the origin and evolution of Earth, the oceans, and life, and maintain and develop the technologies for deep-space return exploration (as demonstrated with Hayabusa), a field in which Japan leads the world.

Expected results and effects

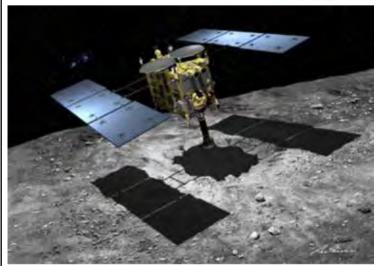
- •By exploring a C-type asteroid, which is rich in water and organic materials, we will clarify interactions between the building blocks of Earth and the evolution of its oceans and life, thereby developing solar system science.
- Japan will further its worldwide lead in this field by taking on the new challenge of obtaining samples from a crater produced by an impacting device.
- •We will establish stable technologies for return exploration of solar-system bodies.

Features:

- •World's first sample return mission to a C-type asteroid.
- World's first attempt at a rendezvous with an asteroid and performance of observation before and after projectile impact from an impactor.
- Comparison with results from Hayabusa will allow deeper understanding of the distribution, origins, and evolution of materials in the solar system.

International positioning:

- Japan is a leader in the field of primitive body exploration, and visiting a type-C asteroid marks a new accomplishment.
- This mission builds on the originality and successes of the Hayabusa mission. In addition to developing planetary science and solar system exploration technologies in Japan, this mission develops new frontiers in exploration of primitive heavenly bodies.
- •NASA too is conducting an asteroid sample return mission, OSIRIS-REx (launch: 2016; asteroid arrival: 2018; Earth return: 2023). We will exchange samples and otherwise promote scientific exchange, and expect further scientific findings through comparison and investigation of the results from both missions.



Hayabusa 2 primary speci(Itatistmetion: Akihiro Ikeshita)

Mass Approx. 609 kg Launch 3 Dec 2014 Mission Asteroid return Arrival 27 June 2018

Earth return 2020

Stay at asteroid Approx. 18 months

Target body Near-Earth asteroid Ryugu

Primary instruments

Sampling mechanism, re-entry capsule, optical cameras, laser range-finder, scientific observation equipment (near-infrared, thermal infrared), impactor, miniature rovers.



Mission flow

Launch Dec 3, 2014



Earth swing-by Dec 3, 2015



Ryugu arrival June 27, 2018



MINERVA-II-1 separation Sep 21, 2018

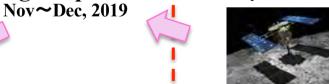


MASCOT separation March 10, 2018

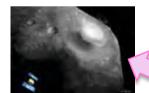


Earth return **End of 2020**





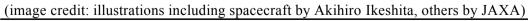
Second touchdown



Impactor (SCI) 5 April, 2019



First touchdown







1. Current project status & schedule overview

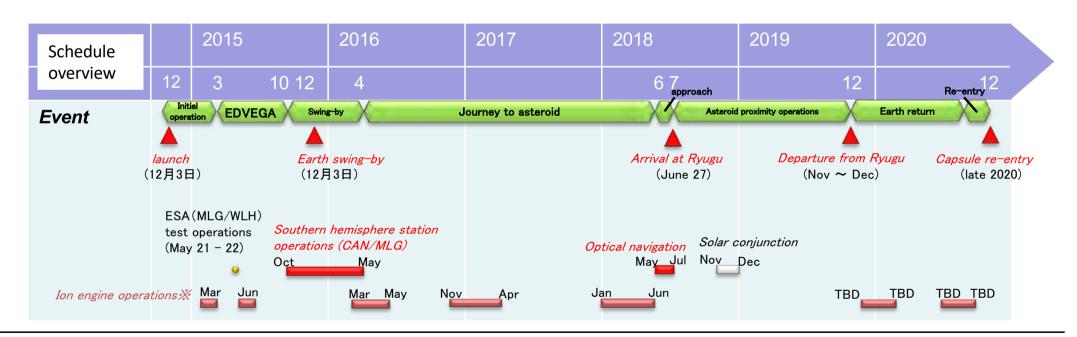


Current

- Implemented the second touchdown from July 9 - 11.

status:

- Touchdown was carried out safely and Hayabusa2 returned to the home position at about 20 km from the center of Ryugu on July 12.
- BOX-C operation is currently underway $(7/20 \sim 31)$. The lowest altitude will be about 5km during 7 / 25~27.







- 2nd touchdown operation: 2019/7/9~11
- Touchdown date & time: 2019/7/11, 10:06:18 JST (on-board time)
- Touchdown location: C01-Cb (Target marker drop area)
- Implemented pinpoint touchdown targeting TM-A dropped during PPTD-TM1A.
- Touchdown was detected through fluctuations in the LRF-S2 ranging value due to deformation of the sampler horn upon touchdown.
- Touchdown position accuracy is 60 cm.



First release

Images from the small monitor camera (CAM-H).
Images before and after touchdown (10x animation)

Capture time:

2019/7/11

Start 10:03:54 (altitude 8.5m) Finish 10:11:44 (altitude 150m)

image interval between 0.5s∼5s



(credit: JAXA)



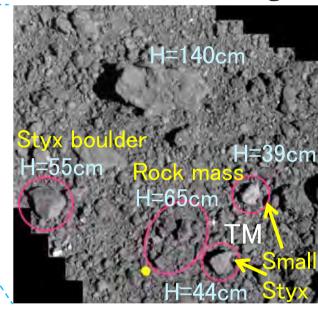


PPTD-TM1 image

CADIT-Cb TM (ravius 3.5m)

TM = target marker (The left-hand image is taken prior to dropping the TM and its position is marked. In the middle image, the TM itself is captured.)

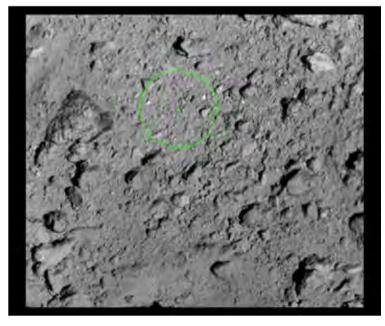
PPTD-TM1B image



H is the maximum estimated height
** boulder names are nicknames, not official designations.

(credit: JAXA、University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST.)

(animation)



DEM (Digital Elevation Map) near the touchdown candidate point

(credit: JAXA、University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST., Kobe University, University of Occupational and Environmental Health)





Challenges for the 2nd touchdown (difference from the 1st)

- ① Due to the optical system on the wide-angle Optical Navigation Camera ONC-W1 becoming cloudy, it was necessary to lower the starting altitude for capturing and tracking the TM (45m to 30m)
 - → For the TM to be in the narrowed field of view of the ONC-W1, the accuracy of the GCP-NAV guidance had to be high.
 - → Managed with the accuracy of the GCP-NAV guidance results
- 2 The TM brightness decreased due to the cloudiness of the optical system of the ONC-W1.
 - → High probability of a bright spot other than the TM being misjudged as the TM
 - → Managed by changing the TM threshold recognition time

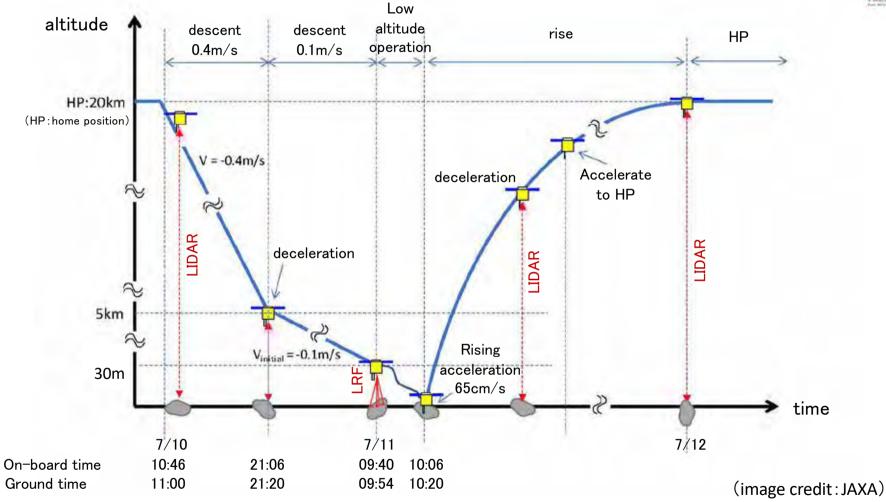
- TM: target marker
 GCP-NAV (Ground Control Point
 Navigation) —method to find the
 - Navigation) →method to find the position and velocity of the spacecraft through observing features on the asteroid surface.
- LRF: Laser Range Finder
- 3 LRF measurable distance decreased due to the cloudiness of the LRF optical system.
 - → Starting altitude for LRF use was lower (17m) than for the first touchdown (28m).

 The descent sequence was therefore shifted to a lower altitude and it was necessary to devise safety measures for the spacecraft.
 - → Timeout was applied.
- 4 Distance measurement error increased due to the cloudiness of the LRF optical system.
 - → Since the range error was predictable, on-board software could correct the range value.
 - → As a result, there was no issue with the LRF range accuracy.



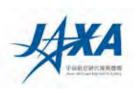


Operation sequence (overall)

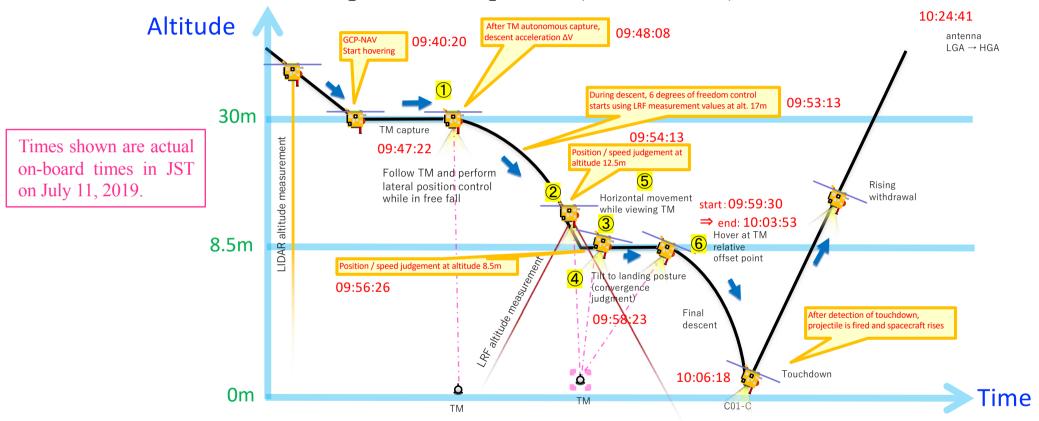


Date & time





Operation sequence (low altitude)



3.0 checkpoints for autonomous judgements as to whether Hayabusa2 continues to the next sequence.

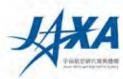
(credit: JAXA)

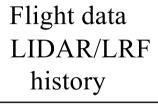


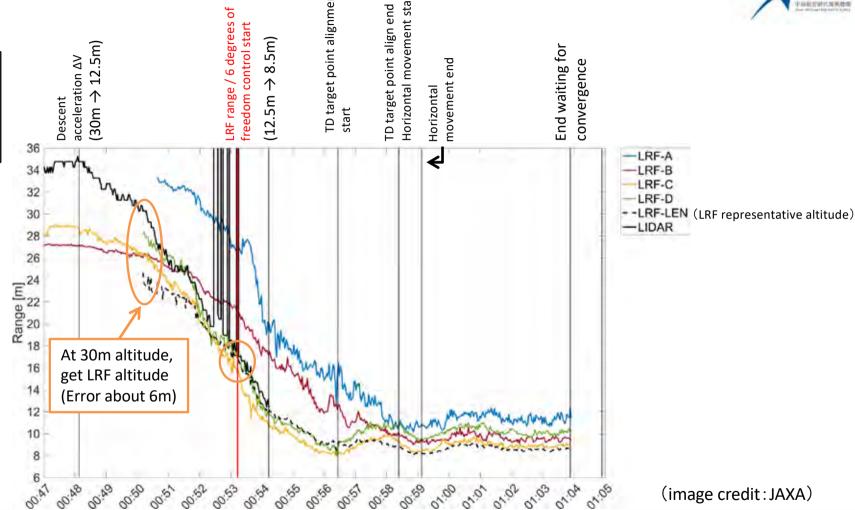


item	Ground time: JST () onboard time		Decision item		
Gate 1	7/10	09:58	Decision made on start of descent (@20km)		
Gate 2	7/10	21:36	End of confirmation on whether to continue descent (@5km)		
Gate 3	7/11	09:04	End of final descent judgement (GO/NOGO judgement)		
HGA→LGA	7/11	10:01 (09:47)	Antenna switching		
Gate 4	7/11	10:01	End of confirmation on switching to LGA		
TD2	7/11	10:20 (10:06)	Touchdown		
LGA→HGA	7/11	10:39 (10:25)	Antenna switching		
Gate 5	7/11	11:10	End check of the state of the spacecraft		
Gate 6	7/11	14:46	Judgement on return to home position		
	7/12	10:50 (10:37)	Return to home position		







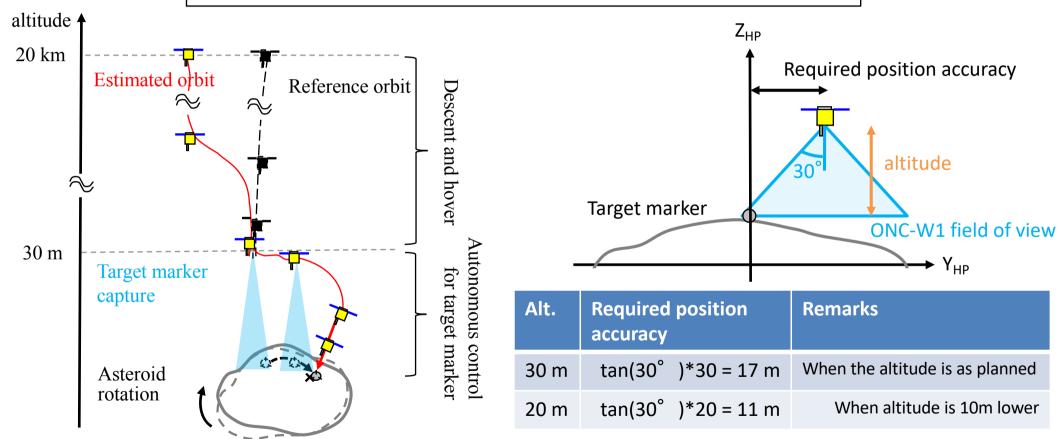


Hayabusa2 reporter briefing





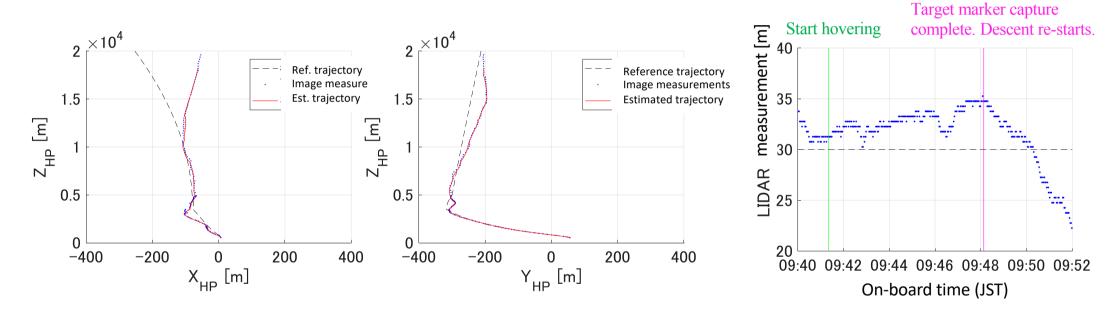
Position accuracy required for target marker capture







Result (descent orbit hovering altitude)



Final position error in the horizontal direction (X, Y, direction) is estimated to be 3m or less, and the final position error in the altitude (Z direction) is 5m or less. Target marker capture was successful.





Countermeasures for decline in camera light reception performance & target marker tracking

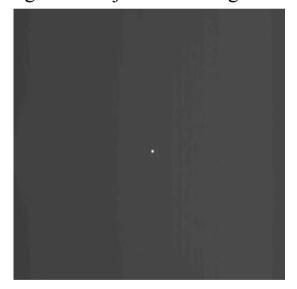
The 1st touchdown in February reduced the light reception performance for the Optical Navigation Camera (ONC-W1).

→ To capture and track the target marker safely and reliably at low altitudes, the image processing parameters (threshold value for digitizing the image into two graduations of black and white) were adjusted.

This step makes it possible to recognize even darker target markers, but also makes it easier to mistake floats (such as dust) around the spacecraft, or bright rocks on the ground, for target markers.

- →Using images acquired during past operations, the perceived motion of a target marker versus floating and similar objects is determined, as well as other identifying parameters (such as the threshold for the time needed to capture the target marker, given the movement between previous and subsequent frames), the size of the target area etc.
- →During the actual mission, the target marker could be tracked stably even in the presence of floats.

Target marker image during tracking at an altitude of 8.5m (Brightness adjusted on the ground)





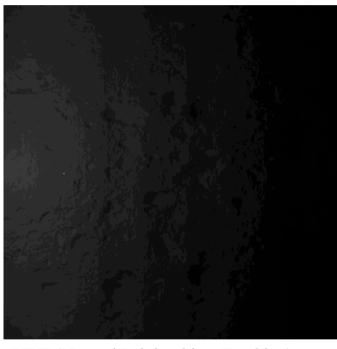


DBT/NBT image and target marker tracking

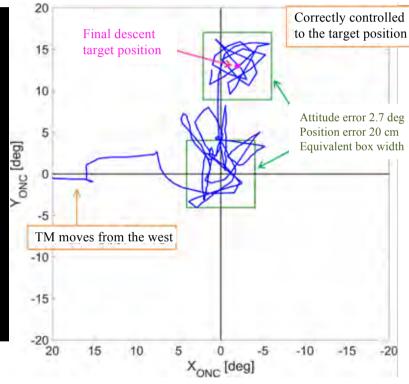


DBT (Differential Bright object Tracking)
: image actually used by the spacecraft for measurement

(Video: 20x)



NBT (Normal Bright object Tracking) : similar image not used for actual measurement (reference)



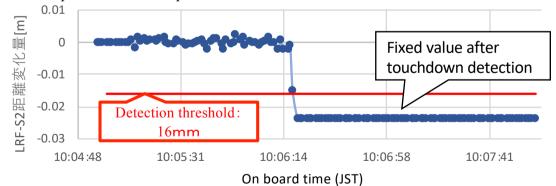
(credit:JAXA)





Final descent below 8.5m & touchdown of the spacecraft

- The final descent ΔV for touchdown was performed at an altitude of 8.5m. Touchdown detection was enabled 50 seconds after the final descent ΔV . Then touchdown occurred on the asteroid surface. Touchdown was judged by detecting the change in distance with the LRF-S2, which measures the distance to the tip of the sampler horn that compresses slightly during touchdown.
- After touchdown detection, the sequence for firing a 3^{rd} projectile, followed by the sampling sequences, were performed. An ascending ΔV then caused the spacecraft to rise and leave the asteroid surface.
- As vibration in the sampler horn is generated by the final descent ΔV , the sequence was devised so as not to generate unnecessary sampler horn vibration.
- At the time of touchdown, the 'tail-up' posture is adopted around the Y-axis to prevent collision of the spacecraft with the boulders and other protrusions, based on the prediction of the spacecraft behaviour.



LRF-S2 measurements before and after touchdown detection

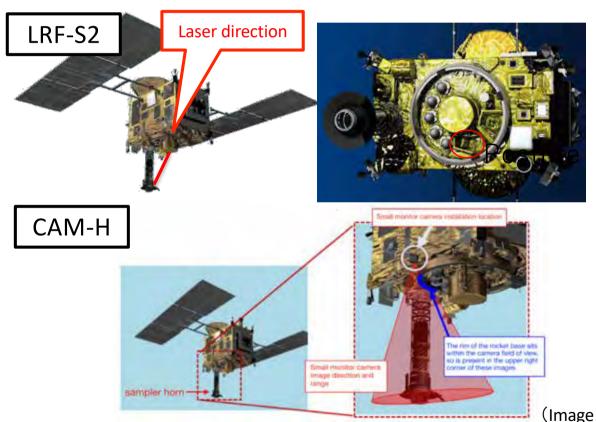
Operation sequence from final descent

Time (s)	On-board time (JST)	Event		
-60		Final altitude control begins		
0	10:04:55	Final descent Δ V begins		
15		RW attitude control beings (sampler horn vibration prevention measures)		
50		Touchdown detection judgement begins		
~70		Posture convergence		
82-84	10:06:17-19	Touchdown detection		
82-84	10:06:17-19	Sampling operation (Projectile launch etc.)		
82-84	10:06:17-19	Rising ΔV begin		
94		RCS attitude control transition		





- The LRF-S2 emits a laser towards a reflector attached to the tip of the sampler horn.
- This measures distance and intensity value.





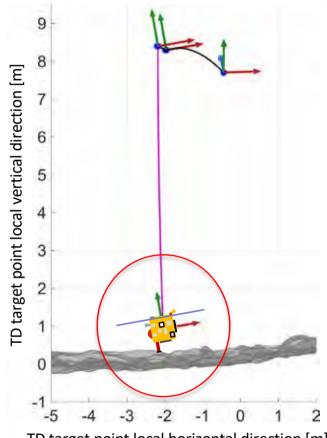
Laser light from the LRF-S2 imaged by CAM-H (At the 2nd touchdown)





Final descent below 8.5m & touchdown of the spacecraft

- During the 2nd touchdown, the tail-up posture was adopted to prevent contact between boulders and other obstructions with the spacecraft.
- During tail-up, in addition to the spacecraft posture aligning to the terrain surface, the posture is rotated by 10 degrees about the Y-axis of the spacecraft to give the target attitude for touchdown.



TD target point local horizontal direction [m]

(Image credit: JAXA)

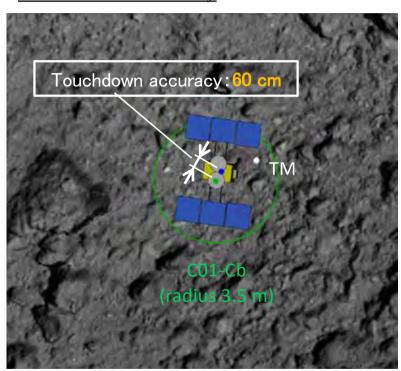
Nominal touchdown attitude during operation plans





2nd touchdown accuracy and sampler horn ground point

2nd touchdown accuracy



Sampler horn ground point



(Credit: JAXA, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST)



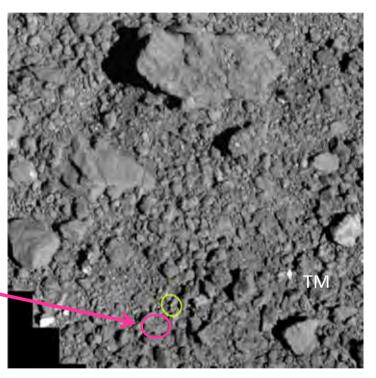


2nd touchdown sampler horn ground point

Sampler horn ground point

CAM-H images





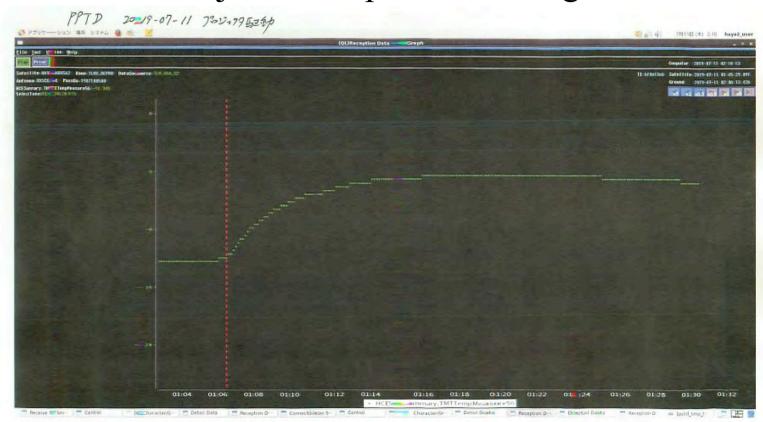
Consistent with CAM-H image

(Credit: JAXA, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST)





Projector temperature change

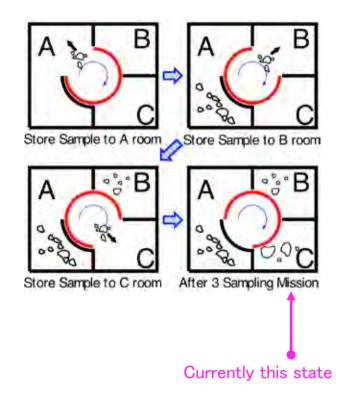






Closing the catcher chamber

- Chamber A closed immediately after the 1st touchdown (February 22).
- Chamber B was open after this, but closed in an operation on June 24. (A total of 7 descent operations were conducted while chamber B was open).
- Chamber C was then open but closed after the 2nd touchdown on July 11 at 14:10 JST (onboard time). (The ascent speed was reduced by 2 cm/s at 13:40 JST so that any sample at the tip of the sampler horn would be collected).





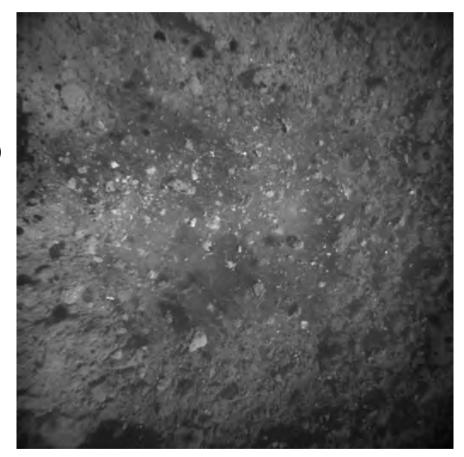


Images from the ONC-W1 Capture time:

2019/7/11

10:06:32 JST (on-board time)

Altitude: about 8m



(Credit: JAXA, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST)





Images from the ONC-W1 Capture time:

2019/7/11

10:08:53 JST (on-board time)

Altitude: about 90m



(Credit: JAXA, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST)



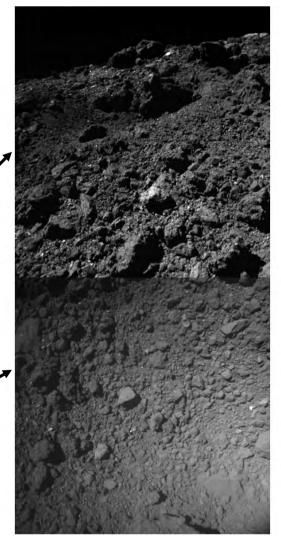


First release

2nd touchdown ONC-W1/W2 composite panoramic image at an altitude of 8m during the final descent.

ONC-W2 2019/07/11 10:04:58 JST (onboard time)

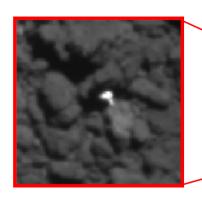
(Credit: JAXA, Chiba Institute of Technology, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Meiji University, University of Aizu, AIST) ONC-W1 2019/07/11 10:04:57 JST (onboard time)





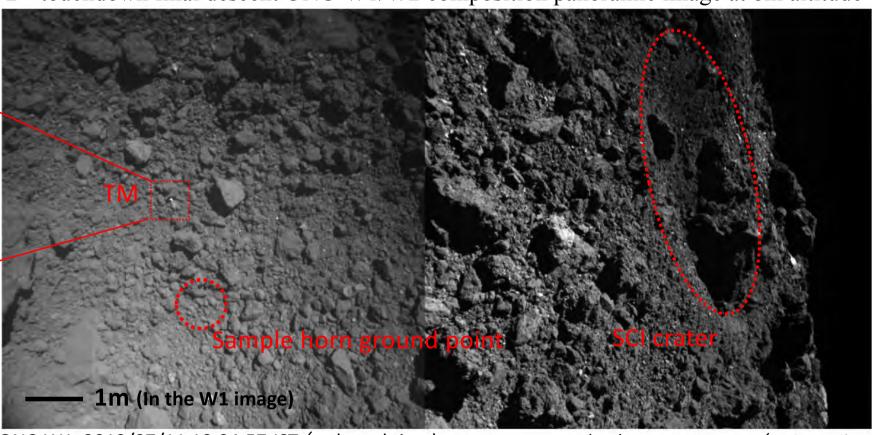


Explanation



(Credit: JAXA, Chiba Institute of Technology, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Meiji University, University of Aizu, AIST)

2nd touchdown final descent ONC-W1/W2 composition panoramic image at 8m altitude



ONC-W1 2019/07/11 10:04:57 JST (onboard time)

ONC-W2 2019/07/11 10:04:58 JST (onboard)



Reference: Geometrical relationship between W1/W2 imaging

- ONC concentrates on TM image acquisition until the final descent at an altitude of 8.5m.
- After the final descent from an altitude of 8.5m, imaging was performed with the W1 and W2 at altitude 8m, 4.7m and 4.2m (planned values).

-20m

W2 field of view

SCI crater

W1 field of view

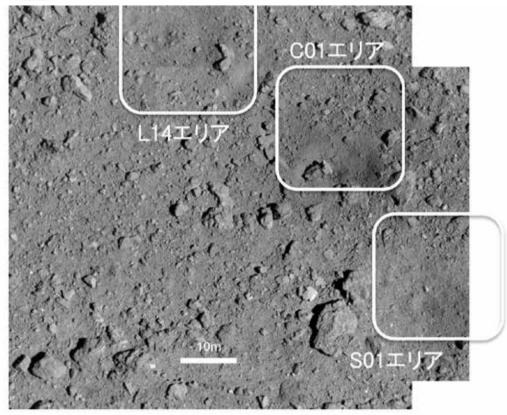
**As the viewing angles for W1 and W2 are slightly larger than 60 degrees, there is a slight overlap.

(Credit: Chiba Institute of Technology)

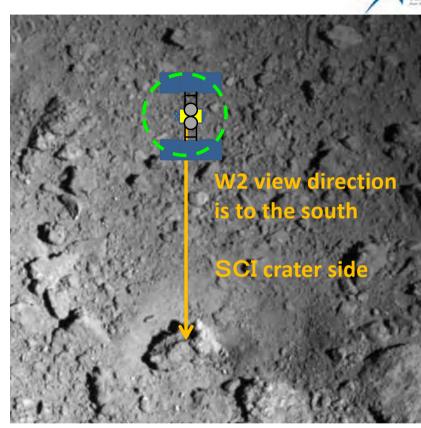


Reference: positional relationship about the C01 area





PPTD—TM1 2019/05/16 Images from an altitude ~0.5km and ~0.6km



PPTD-TM1

(Credit: JAXA, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST)



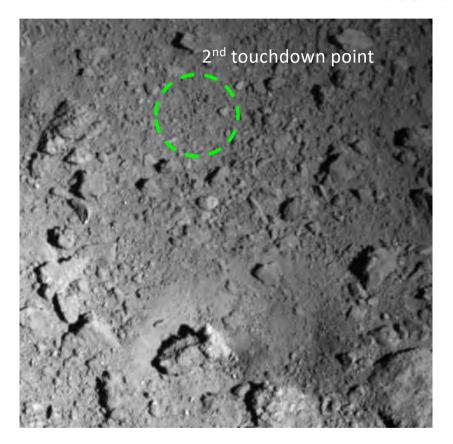
4. Name of the 2nd touchdown point



The name of the 2nd touchdown point is:

Uchide-no-kozuchi

Meaning: In Japanese folklore, the uchide-no-kozuchi is a magic hammer that can produce great riches. The samples gathered from this site are expected to produce great scientific results.



(Image credit: JAXA, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST)



5. Upcoming events



- Operation plans
 - BOX-C operation from July 20 31. The lowest altitude will be about 5km from July 25 27.
- Press and media briefings
 - 8/22 (Thursday) 15:00~16:00 regular press briefings @ Tokyo office



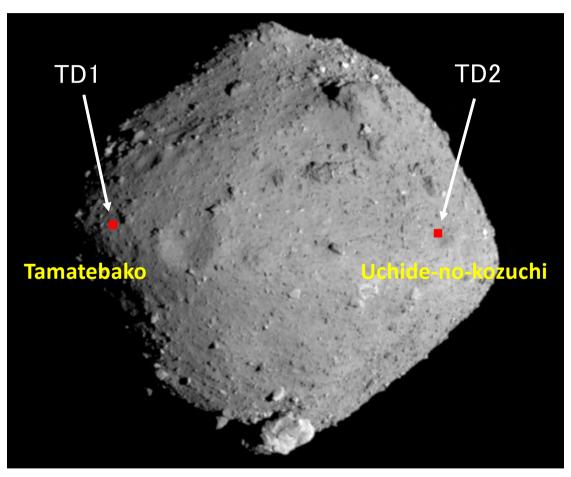


Reference material



Locations for the 1st (TD1) and 2nd (TD2) touchdown





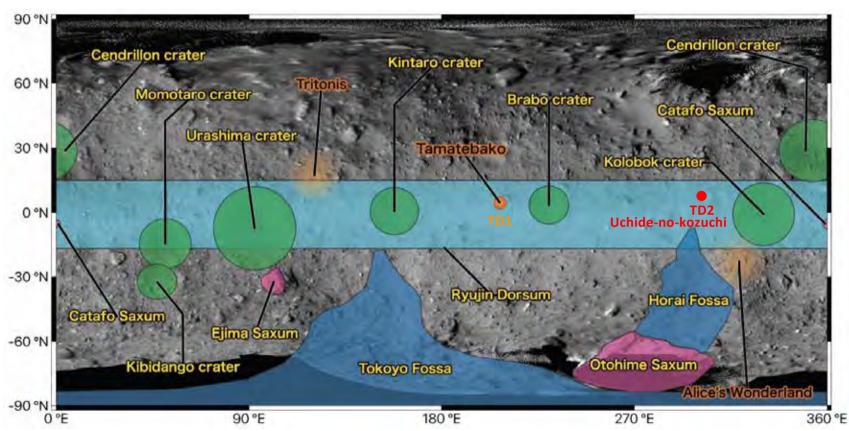
2019/5/20 Taken from the home position

(Image credit: JAXA, University of Tokyo, Kochi University, Rikkyo University, Nagoya University, Chiba Institute of Technology, Meiji University, University of Aizu, AIST)



Touchdown locations for the 1st (TD1) and 2nd (TD2) touchdown





Note: Tritonis (landing site for MINERVA-II1), Alice's Wonderland (MASCOT landing site), Tamatebako (first touchdown point) are nicknames and not recognised by the International Astronomical Union (IAU). Other places names are official names recognised by the IAU.



Optical navigation camera (ONC)

NAGOYA

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ONC: Optical Navigation Camera

Objective: Images fixed stars and the target asteroid for spacecraft guidance and scientific measurements

Scientific measurements:

- Form and motion of the asteroid:

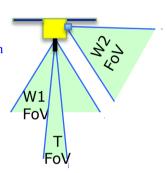
 Diameter, volume, direction of inertial principal axis, nutation
- Global observations of surface topography
 Craters, structural topography, rubble, regolith distribution
- Global observations of spectroscopic properties of surface materials

Hydrous mineral distribution, distribution of organic matter, degree of space weathering

High-resolution imaging near the sampling point
 Size, form, degree of bonding, and heterogeneity of surface particles; observation of sampler projectiles and surface markings



- Elucidation of features of target asteroid
- Distribution of hydrous minerals and organic matter, space weathering, boulders
- Sampling site selection
- Basic information on where to collect asteroid samples
- · Ascertaining sample state
- High-resolution imaging of sampling sites





	ONC-T	ONC	-W1	ONC-W2
Detector	20	× 1024 px)		
Viewing direction	Downward (telephoto)		Downward (wide- angle)	Sideward (wide-angle)
Viewing angle	6.35° × 6.35°	65.24° × 65.24°		
Focal length	100 m-∞	1 m-∞		
Spatial resolution	1 m/px @ 10-km alt. 1 cm/px @100-m alt.	10 m/px @10-km alt. 1 mm/px @1-m alt.		
Observation wavelength	390, 480, 550, 700, 860, 950, 589.5	485–655 nm		

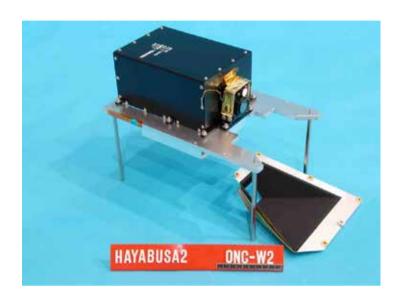
(© JAXA)



ONC-W2 mounting position







- Mounted on the side. Diagonal-downwards imaging possible.
 - Earth imaging during swing-by
 - MASCOT separation imaging
 - SCI crater search operation on Ryugu