



Candidates for landing sites for the Hayabusa2 mission

August 23, 2018

JAXA Hayabusa2 Project







Regarding Hayabusa2:

 Candidates for landing sites for touchdown, MASCOT, and MINERVA-II



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Overview of Hayabusa2



Objective

We will explore and sample the C-type asteroid Ryugu, which is a more primitive type than the S-type asteroid Itokawa that Hayabusa explored, and elucidate interactions between minerals, water, and organic matter in the primitive solar system. By doing so, we will learn about the origin and evolution of Earth, the oceans, and life, and maintain and develop the technologies for deep-space return exploration (as demonstrated with Hayabusa), a field in which Japan leads the world.

Expected results and effects

- By exploring a C-type asteroid, which is rich in water and organic materials, we will clarify interactions between the building blocks of Earth and the evolution of its oceans and life, thereby developing solar system science.
- Japan will further its worldwide lead in this field by taking on the new challenge of obtaining samples from a crater produced by an impacting device.
- •We will establish stable technologies for return exploration of solar-system bodies.

Features:

- •World's first sample return mission to a C-type asteroid.
- World's first attempt at a rendezvous with an asteroid and performance of observation before and after projectile impact from an impactor.
- Comparison with results from Hayabusa will allow deeper understanding of the distribution, origins, and evolution of materials in the solar system.

International positioning:

- Japan is a leader in the field of primitive body exploration, and visiting a type-C asteroid marks a new accomplishment.
- This mission builds on the originality and successes of the Hayabusa mission. In addition to developing planetary science and solar system exploration technologies in Japan, this mission develops new frontiers in exploration of primitive heavenly bodies.
- NASA too is conducting an asteroid sample return mission, OSIRIS-REx (launch: 2016; asteroid arrival: 2018; Earth return: 2023). We will exchange samples and otherwise promote scientific exchange, and expect further scientific findings through comparison and investigation of the results from both missions.



(Illustration: Akihiro Ikeshita)

Hayabusa 2 primary specifications

Mass	Approx. 609 kg
Launch	3 Dec 2014
Mission	Asteroid return
Arrival	27 June 2018
Earth return	2020
Stay at asteroid Target body	Approx. 18 months Near-Earth asteroid Ryugu

Primary instruments

Sampling mechanism, re-entry capsule, optical cameras, laser range-finder, scientific observation equipment (near-infrared, thermal infrared), impactor, miniature rovers.



Mission flow





subsurface samples

(Illustrations: Akihiro Ikeshita)

asteroid's surface



1. Current project status & schedule overview



Current status:

- On August 5, we began operations for gravity measurement, approaching Ryugu to an altitude minimum of 851m at around 08:10 JST on August 7. The spacecraft then returned to the home position on August 10.
- Based on the data obtained to date, we investigated possible landing points.
- BOX-B operations began on August 18 (scheduled to return to BOX-A on September 7).

Schedule overview:





Points to note in this document



Pay attention to the orientation of the asteroid!

Public images so far have shown the northern direction of the Solar System (direction of the Earth's North Pole) pointing upwards.



(©JAXA, U. of Aizu et al.)

In this document, like a typical map, north is drawn pointing upwards. As Ryugu rotates backwards, Ryugu's "north" is at the bottom of the left-hand figure, but the top in this map (top and bottom are reversed). From now on, public images also show the northern direction of Ryugu pointing upwards.



Ryugu has a reverse rotation (retrograde)

Example of a Ryugu "map"





Landing Site Selection (LSS) conference:

- •Held on August 17, 2018
- 109 attendees (including remote participants)
- •39 participants from overseas (14 from DLR, 2 from CNES and 2 from NASA)
- Discussion took place from 10 am to 7 pm.

Determined candidate landing points for spacecraft touchdown, MASCOT & MINERVA-II







Determined landing site candidates Touchdown : L08(backup:L07,M04) MASCOT : MA-9 MINERVA-II-1 : N6







Determined landing site candidates

- Touchdown : L08(backup:L07, M04)
- MASCOT : MA-9

MINERVA-II-1 : N6



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Operation Schedule

Touchdown 1 rehearsal 1:September 11 ~ 12
(Arrival at lowest altitude : September 12)MINERVA-II-1 operation:September 20 ~ 21
(MINERVA-II-1 separation : September 21)MASCOT operation:October 2 ~ 4
(MASCOT separation : October 3)

Touchdown 1 rehearsal 2:mid-OctoberTouchdown 1:late-October

Note: date of operations may be changed.



3. Selection of touchdown site candidates

Prerequisites for touchdown candidate points, part 1

Range on the celestial body where the spacecraft can land

The spacecraft can move along the line connecting the Earth and Ryugu.

The region where the spacecraft can land is the area within about 200 m to the north and south of Ryugu's equator (range of about $\pm 30^{\circ}$ latitude)





3. Selection of touchdown site candidates

Prerequisites for touchdown candidate points, part 2

Surface conditions suitable for landing

(1) Average slope within $30^{\circ} \times \leftarrow$ Limit due to orientation of solar panels

(2) Flat region with 100m diameter \leftarrow Navigation guidance accuracy

(3) Boulder height less than 50cm \leftarrow Length of sampler horn

(4) Absolute temperature less than 370K(97°C)

 \leftarrow Within operating temperature range of equipment





Procedure of the selection of landing site candidates





Narrow down touchdown candidates: Step 1, from shape model





Narrow down touchdown candidates: Step 1, from shape model





3. Selection of touchdown site candidates

Narrow down touchdown candidates: Step 2, from images



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Narrow down touchdown candidates: Step 2, from images

Candidate sites: four low latitude ("L") and three mid-latitude ("M")



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3. Selection of touchdown site candidates

Narrow down touchdown candidates: Step 3, feasibility

Further narrowing down from 7 candidates (4 low latitude, 3 mid-latitude)





Narrow down touchdown candidates: Step 3, feasibility

Further narrowing down from 7 candidates (4 low latitude, 3 mid-latitude) Create boulder map and check boulder coverage



(©JAXA、U. Tokyo et al)

Example boulder map



3. Selection of touchdown site candidates

L08

Narrow down touchdown candidates: Step 3, feasibility

Further narrowing down from 7 candidates (4 low latitude, 3 mid-latitude)

Creating boulder maps







Boulder area (D > 3 m) : 16.8%





Boulder area (D > 3 m) : 12.7%

Distribution of boulders. The size of boulders with a diameter greater than 3m is indicated by marker size and color. Brown is for boulders larger than 10m.

(©JAXA)

L12 Boulder area (D > 3 m) : 17.6%



3. Selection of touchdown site candidates

Narrow down touchdown candidates: Step 3, feasibility

Further narrowing down from 7 candidates (4 low latitude, 3 mid-latitude)

Determining boulder coverage







Features of Ryugu

- Top shape with a very circular equatorial bulge
- Radius: mean ~450 m (equatorial ~500 m, polar ~440 m)
- Mass: ~ 450 million ton $(GM \sim 30 \text{ m}^3 \text{s}^{-2})^{\text{*}}$
- Rotation axis: $(\lambda, \beta) = (180^\circ, -87^\circ)$
- Obliquity: $\sim 8^{\circ}$
- Rotation period : P = 7.63 hours
- Reflectance factor (v-band) : 0.02
- Crater number density: as much as those on Itokawa and Eros
- Many boulders: the largest near the south pole is ~ 130 m across
- Optical spectra: flat spectra, bluer in equatorial bulge and poles
- NIR spectra: uniform flat (slightly redder) spectra with weak water absorption
- brightness temperature : strong roughness effect (flat diurnal Temperature variation), higher thermal inertia in the equatorial bulge

(※The gravity at the equator is eighty-thousandth of the Earth and a few times of Itokawa)









Scientific evaluation points of landing site candidates

- Point 1: surface properties
 - •Examine surface temperature and thermophysical properties using the Thermal Infrared Imager data.
 - •Examine composition differences using spectral data from the Near Infrared Spectrometer.
 - •Examine terrain, geology and space weathering from different wavelength images from the Optical Navigation Camera.
 - * Evaluate potential scientific merit.
- Point 2: safety
 - •Examine size and spatial distribution of boulders (from images and laser altimeter data).
 - * Evaluate safety by estimating the number of small boulders from the boulder size distribution.
- Point 3: sample yield
 - •Estimate surface grain sizes from observational data from the Thermal Infrared Imager. \downarrow
 - * From the particle size of surface regolith, evaluate where the maximum amount of sample can be gathered.





The points for scientific evaluation of the landing site candidates: one example from observation data.

ONC-T spectral slope map (Box-C)



(©JAXA, University of Tokyo & collaborators)





Summary of the scientific evaluation of landing point candidates

	Candidate site	Point 1 surface properties	Point 2 safety	Point 3 Sample yield	Total	
	L5 L7 L8 L12 M1	22 22 22 22 22 21	29 31 31 31 31 33	12 12 12 11 13	63 65 65 64 67	Although no big difference in evaluation, L08 & L07 are good for low latitude (L). M1 & M4
	M3 M4	21 21 21	30 33	13 13	64 67	are good if mid-latitude (M) is preferred.
Su in bi w th Sp th m	urface temp frared spec g differenc hich is alm ere are min pectral chan at lower lat ore diverse	berature, visible / ne etrum etc. There is n es across the surfac ost uniform. Howev for differences. cacteristics indicate titudes may contain particles.	ear- no ver, Ver, Safety numbe smaller	Mid-latitudes (M) sized particles, wh sample yield. is evaluated from r density of boulders boulders and surfac	seem to ich is ex informa s, estimat e roughn	b have many smaller pected to increase the ation such as ted number of ness.





Summary of the scientific evaluation of landing point candidates

Scientifically important points:

- (1) Difference in mixing ratio of Ryugu surface material has a low dependence on location.
- (2) A diversity of different materials are present mixed together on Ryugu's surface.
- (3) Regardless of where the sample is collected, there is a high possibility of gathering diversified materials that represent the whole of Ryugu.





MASCOT (Mobile Asteroid Surface Scout)

- Created by DLR (German Aerospace Center) and CNES (French National Centre for Space Studies)
- Small lander with mass approx. 10 kg
- Carries four scientific instruments
- Can move only once, by jumping

Scientific instruments aboard MASCOT

Device	Function
Wide-angle camera (MASCAM)	Imaging at multiple wavelengths
Spectroscopic microscope (MicrOmega)	Investigation of mineral composition and characteristics
Thermal radiometer (MARA)	Surface temperature measurements
Magnetometer (MASMAG)	Magnetic field measurements





Flight model (© DLR)



MASCOT







Criteria for the selection of landing site candidates for MASCOT

- Probability to land outside the candidates for touchdown > 95 %
- Probability to have RF link during at least 40% of the asteroid period > 90 %
- Probability to have a good daylight ratio (between 40% and 70%) > 90 %
- Selection among remaining candidates to have one candidate per reachable and suitable zone, in northern as well as in southern hemisphere
- No overlap with the landing site candidates for MINERVA-II

computed 10,000 ~ 100,000 Monte-Carlo trials



Further conditions

- Temperature (illumination, orientation, heating, prior separation, asteroid surface temperature, night temperature underneath MASCOT)
- •Operation
- •Boulders
- Observation conditions for four instruments





Scientific Criteria for Landing Site

Instrument	Criteria for Landing Site
MicrOmega	Composition of C-rich and/or OH-rich content. Low Temperature of landing site for high SNR (signal-to-noise ratio).
MASCAM	Boulders in the field of view Fine grained particles Compositional heterogeneity (color!) Fresh material (non-weathered)
MARA	Thin to no regolith layer. Landing or hopping close to boulders preferable. Preferentially less thermally altered region.
MASMAG	Low thermal inertia for fine grained fraction of the regolith Homogeneous thermal inertia (link the MARA + TIR data) Low rock abundance -> no inhomogeneities in the FoV (field of view)



Remote sensing instruments onboard HY2 give preliminary information on Ryugu : no big variation over the surface \rightarrow homogenous





Landing site candidates for MASCOT: selected 10 candidates (MA-1~MA-10)



^{(©}DLR,CNES)

Light blue shows the first contact points and blue shows the first settlement points. The candidates of touchdown sites are shown by red squares (low latitude) and pink squares (mid-latitude).





MASCOT Ranking meeting (Toulouse Aug 14)



Fruitful meeting, useful rehearsal and training (thanks to JAXA) => complex process but all milestones/deliveries were reached according to the schedule





Landing site candidates for MASCOT: Selection from 10 candidates and the order of priority



Priority of landing site: MA-9 > MA-1 > MA-10 > MA-7 > MA-5 > MA-2



5. Selection of landing site candidates for MASCOT MASCOT landing site







6. Selection of landing site candidates for MINERVA-II



Release of MINERVA-II-1 A & B (MINERVA-II-2 will be released next year)



MINERVA-II-2

Created by the MINERVA-II consortium (Tohoku University, Tokyo Denki University, Osaka University, Yamagata University, Tokyo University of Science)



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MINERVA-II-1A, 1B



Produced at JAXA

MINERVA-II-1A, MINERVA-II-1B Specification

- •size:diameter 17cm, height 7cm
- •weight:about 1kg each
- •Actuator: 2 DC motors
- Mounted sensor : camera, photodiode, accelerometer, thermometer, gyro.
- Communication speed: 32kbps(max)



6. Selection of landing site candidates for MINERVA-II



Conditions for MINERVA-II landing site selection:

- Landing site does not overlap with spacecraft touchdown candidates.
- Landing site does not overlap with MASCOT landing site candidates.
- The altitude of the spacecraft after separation must not be lower than 30m.
- Ensure communication with ground station.
- Not high temperature region, and fewer parts in shadow
- Due to the equatorial ridge, separation near the equator results in widely spaced landing points to the north and south.
- Separating in the southern hemisphere may result in a spacecraft altitude below 30m.

• Separate in northern hemisphere, more than 100m north of the equator.





6. Selection of landing site candidates for MINERVA-II



Landing site candidates for MINERVA-II: northern hemisphere



- •Touchdown•confirm no overlap with MASCOT's landing site.
- •Also consider observability etc. using the ONC-T camera. -

Candidate locations: N6 > N1 > N7







Detail of surface from low altitude



How far can the navigation guidance accuracy be increased?

Ryugu surface taken from an altitude of about 1km.



8. Future Plans



■Schedule for press briefings

- Sept. 5 (Wed) 11:00~12:00
- Sept. 27 (Thurs) 14:30~15:30

■Outreach and events (in Japanese)

- Events for Children
 - Why Hayabusa2? Any questions classroom
 - Sept. 2 (Sunday) 2 4pm
 - Sagamihara City Museum
 - Online broadcast planned